Measurements of Neutrino-Electron Scattering Cross-Section with CsI(TI) Scintillating Crystal Detector at the Kuo-Sheng Reactor Neutrino Laboratory

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# OUTLINE

- TEXONO Physics Program
- $\overline{v}_{e} e^{-}$  Scattering Motivation
- TEXONO Experiment CsI(TI) Array
- Event Selection & Data Analysis Outline
- Background Understanding & Suppression
- Preliminary Results
- Plans & Summary

# **TEXONO** Physics Program

- TEXONO Collaboration : Taiwan(AS, INER, KSNPS, NTU), China(IHEP, CIAE, THU, NJU), Turkey(METU), India(BHU)
- Program : Low Energy Neutrino & Astroparticle Physics



- [1] Magnetic Moment Search at  $\sim 10 \text{ keV} \rightarrow \text{PRL 2003}$ , PRD 2007
- [2]  $\sin^2 \theta_W$  measurement at ~ MeV range  $\rightarrow$  This Talk Preliminary Results [3]  $v_e N$  Coherent Scattering & WIMP Search at sub keV range

# $\mathbf{\bar{v}}_{e}$ – $e^{\scriptscriptstyle -}$ Scattering



 $(R_{CC}:R_{NC}:R_{Int})$ 

 $R_{SM}(\nu_{e}e) \rightarrow (1.83:0.17:0.99)$ 

A basic SM process with CC, NC & Interference

Not well-studied in reactor energy range ~
MeV

### Experimental Approach -- CsI(TI) Crystal Scintillator Array:

- $\textcircled{\mbox{o}}$  proton free target (suppress  $\overline{v_e}\mbox{-}p$  background)
- scale to  $\vartheta$  (tons) design possible
- good energy resolution, alpha & gamma
   PSD
- allows measure energy, position, multiplicity

more information for background understanding & suppression

• focus at >3 MeV recoil energy range  $\rightarrow$  less ambient background & reactor  $\nabla_{e}$  spectra well known.

$$\delta[\sin^2 \theta_{\rm W}] \sim \left\{ \begin{array}{c} 0.14 \cdot \ \delta[\xi(\bar{\nu}_{\rm e}e)] \\ 0.32 \cdot \ \delta[\xi(\nu_{\rm e}e)] \end{array} \right\} \xi = \frac{R_{expt}(\nu)}{R_{\rm SM}(\nu)}$$

### Kou-Sheng Reactor Power Plant



Kuo-Sheng Nuclear Power Station : Reactor Building



### KS NPS -II : 2 cores × 2.9 GW



### Total flux about 6.4x10<sup>12</sup> cm<sup>-2</sup>s<sup>-1</sup>

KS v Lab: 28m from core #1

### 30 mwe overburden

## **Kuo-Sheng Reactor Neutrino Laboratory**



### **Inner Target Volume & Shielding**



## CsI (TI) Array : Basic Performance



### **Event Selection**

**Reactor OFF**  $\underset{0}{\mathsf{R}_{\mathsf{exp}}}_{\mathsf{day}}^{\mathsf{day}^{-1}} \mathsf{kg}^{-1} [\mathsf{10keV}]^{-1}$ 4 Ω  $\mathbb{R}_{exp}^{exp}(day^{-1}kg^{-1}[10keV]^{-1})$ ····· CRV-Tag All Vetos <sup>137</sup>Cs <sup>63</sup>Cu 2 208-10 13 5 7 9 15 3 11 Energy (MeV) 1 -1 10 --- Raw Data --- CRV -2 10 ..... MHV ---- PSD -3 Z-Pos Cut 10 -4 102000 3000 4000 5000 1000 6000 7000 8000 0 Energy (keV)

CUTS (3 - 8 MeV)	Efficiencies DAQ Live Time Eff. 88.78%		
CRV	92.7 % 99.9 % ~100 % 80% 77.1 %		
MHV			
PSD			
Z-pos			
Total			

$$\frac{S}{B} \cong \frac{1}{15}$$
 at 3 MeV

# Background Understanding: Due to <sup>208</sup>TI and Cosmic I nefficiency



# Background Prediction via PAIR PRODUCTION



### **Residual Background Understanding & Suppression** Idea -- Use Multiple Crystal Hit (MH) spectra to predict Single Crystal Hit (SH) Background to the neutrino events Two Background Sources : Cosmic Rays and <sup>208</sup>TI $\frac{MH_{non\ cos}}{MH_{tot}})_{ON,OFF} = 1 - \varepsilon = \left(\frac{SH\left[\nu BKG\left(\cos\right)\right]}{SH_{tot}}\right)_{ON,OFF}$ **Cosmic Ray** (3-8 MeV) SH[vBKG(2614 + 583)] SH[2614 + 583(MC)]208**TI** MH[2614;583(data)] MH[2614;583(MC)](3-3.5 MeV) ε<sub>CRV</sub> ~ 93 % **BKG(SH)** Sources combined **BKG(SH)** from three measurements: Direct Reactor OFF(SH) spectra Energy (MeV) Tl-208/Cu cosmic Predicted BKG(SH) from OFF(MH) ~ 35% Tl 3.0 - 3.5~ 65% (γ,γ) Predicted BKG(SH) from ON(MH) 3.5 - 6.5~ 100%

v = ON(SH) - BKG(SH)

Studies of Neutrino-Electron Scattering

6.5 - 8.0

~ 45% Cu

 $(\mathbf{n}, \boldsymbol{\gamma})$ 

~ 55%

#### **Cross Section & Weinberg Angle** Cross-Section: $R = [1.18 \pm 0.29 (stat) \pm 0.08 (sys)] \times R_{sm}$ $\sin^2 \theta_w = 0.264 \pm 0.040 (stat) \pm 0.010 (sys)$ Weinberg Angle: 0.015 9A $\chi^2 / dof = 9.9/9$ SM TEXONO 0.0125 Best-Fit 0.4 $R_{expt}$ (day<sup>-1</sup>kg<sup>-1</sup>[500keV]<sup>-1</sup>) Sys. Error 0.01 0.2 0.0075 0.005 gv 0.4 -0.2 0.2 -0.4 0.0025 \_ I¶ I III нŦн -0.2 0 н н LSND -0.0025Sin<sup>2</sup>θ<sub>w</sub> -0.0050 0.5 -0.00756.5 7 7.5 3 3.5 5.5Energy (MeV)

# Summary Table with Other Experiments

	Experiment	Energy (MeV)	Events	<b>Cross-Section</b>	sin²θ <sub>w</sub>
v <sub>e</sub> -e-{	LAMPF [Liquid Scin.]	7 - 60	236	$[10.0 \pm 1.5 \pm 0.9]$ x E <sub>ve</sub> 10 <sup>-45</sup> cm <sup>2</sup>	0.249 ± 0.063
	<b>LSND</b> [Liquid Scin.]	10 - 50	191	$[10.1 \pm 1.1 \pm 1.0] \\ x E_{ve} 10^{-45} cm^2$	0.248 ± 0.051
v <sub>e</sub> −e-	Savannah-River [Plastic Scin.]	1.5 - 3.0 3.0 – 4.5	381 71	[0.86 ± 0.25] x σ <sub>V-A</sub> [1.70 ± 0.44] x σ <sub>V-A</sub>	0.29 ± 0.05
	Savannah-River Re-analysed (PRD1989, Engel&Vogel)	1.5 – 3.0 3.0 – 4.5	N/A	[1.35 ± 0.4] x $\sigma_{SM}$ [2.0 ± 0.5] x $\sigma_{SM}$	N/A
	Krasnoyarsk (Fluorocarbon)	3.15 – 5.18	N/A	[4.5 ± 2.4] x 10 <sup>-46</sup> cm <sup>2</sup> /fission	0.22 ± 0.75
	Rovno [Si(Li)]	0.6 – 2.0	41	[1.26 ± 0.62] x 10 <sup>-44</sup> cm <sup>2</sup> /fission	N/A
	MUNU [CF <sub>4</sub> (gas)]	0.7 – 2.0	68	1.07 ± 0.34 events day <sup>-1</sup>	N/A
	TEXONO [CsI(TI) Scin.]	3 - 8	~ 450	[1.18 ± 0.29 ± 0.08] x R <sub>SM</sub>	0.264 ± 0.042

### Study of Systematic Uncertainties Approach – Use non-v events for demonstration

#### <sup>208</sup>TI Peak Events 4 Count kg<sup>-1</sup>day<sup>-1</sup> 3.75 3.5 3.25 3 2.75 2.5 2.25 2 25 50 75 100 150 175 200 0 125 Time (days)

### <sup>208</sup>TI (SH) Prediction



### BKG – Pred.



#### ON-OFF Stability < ~0.5%

- Random trigger events for DAQ & Selection Cuts
- **Stability** of **TI-208 (2614 keV)** peak events

Cosmic Induced BKG(SH) Prediction < ~1.2%

- Successfully **Predict Cosmic BKG** at **NEUTRINO FREE REGION**
- TI-208 Induced BKG(SH) Prediction <~3%
- Successfully Predict TI-208 Induced BKG(SH) >3MeV at Reactor OFF periods

Successfully **Predict TI-208** peak intensity for both

Reactor **ON/OFF** with the same tools (MC)

### Interference Term



# Neutrino Magnetic Moment and Charge Radius

**Neutrino Magnetic Moment** 

$$R(ON - BKG) = R(SM) + \kappa^{2} \times R(MM \ [\mu_{v} = 10^{-10} \ \mu_{B}])$$

$$\kappa^{2} = 0.9 \pm 2.51(\text{stat}) \pm 0.71(\text{sys})$$

$$\mu_{v} < 2.4 \times 10^{-10} \times \mu_{B}$$
at 90 % C. L.

### **Neutrino Charge Radius**

$$\sin^2 \theta_W \rightarrow \sin^2 \theta_W + (\sqrt{2}\pi\alpha / 3G_F) \left\langle r_{\overline{\nu}_e}^2 \right\rangle$$

$$\left\langle r_{\bar{\nu}_{e}}^{2} \right\rangle = \left[ 1.23 \pm 1.71 (stat) \pm 0.36 (sys) \right] \times 10^{-32} \ cm^{2}$$

 $\chi^2 / dof = 9.8/9$ 

## Summary and Status

- 200 kg CsI(TI) Scintillating Crystal Array Analysis Threshold at 3 MeV
- Preliminary Results:
  - $\sigma(\bar{\nu_e} e^-)$  with  $\sim 25\%$  accuracy
  - **Weinberg Angle** with  $\sim 15\%$  accuracy
  - verify SM negative interference
  - $\mu_{\nu}$  sensitivity  $\sim 10^{-10}$
  - $\square$  neutrino charge radius sensitivity  $\sim 10^{-32}$